

THE NUMEROV ALGORITHM FOR
MAGNETIC FIELD TRAJECTORIES
by
Peter Signell

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## Input Skills:

1. Use the Taylor series to expand a function about a point (MISN-0-4).

## Output Skills (Knowledge):

K1. Derive the recurrence relation for the Numerov Algorithm, to second order and in two dimensions, in a form suitable for use in obtaining the trajectory of a charged particle in an arbitrary magnetic field. Show all steps in the derivation.
K2. Derive equations for insertion of initial position and velocity in the Numerov Algorithm and communicate a method of obtaining a particular desired accuracy.

## Post-Options:

1. "Trajectory of a Charged Particle in a Magnetic Field: Cyclotron Orbits (a computer project)" (MISN-0-127).

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## 1. Introduction and Description

Many problems in science and engineering can not be solved in terms of known functions, even when the under lying equation is known. Such a problem is the trajectory of a charged particle in a non-uniform magnetic field. For such cases one must resort to general numerical techniques: one of the most common is examined in this module.

## 2. Study Material

The force on a charged particle in a magnetic field is the Lorentz force:

$$
\begin{equation*}
\vec{F}=q \vec{v} \times \vec{B} \tag{1}
\end{equation*}
$$

where $q$ is the charge of the particle, $\vec{v}$ is its velocity, and $\vec{B}$ is the value of the magnetic field at the present location of the particle. The present force on the particle is $\vec{F}$. The force influences the particle's trajectory through Newton's Second Law:

$$
\begin{equation*}
\vec{F}=m \vec{a} \tag{2}
\end{equation*}
$$

where $m$ is the particle's mass.
In our case $\vec{B}$ will always be at right angles to $\vec{v}$, as is obvious from Eq. (1), hence $\vec{v}$ can change $\vec{B}$ 's direction but not its magnitude. We will restrict ourselves to motion in the $x-y$ plane by putting $\vec{v}$ there initially and putting $\vec{B}$ in the $z$-direction:

$$
\begin{align*}
& \vec{B}=B(x, y) \hat{z} \\
& \vec{v}=x^{\prime} \hat{x}+y^{\prime} \hat{y} \tag{3}
\end{align*}
$$

where a prime denotes derivative with respect to time: $x^{\prime} \equiv d x / d t$.


Figure 1. A function $x(t)$, specified at equally spaced values of $t$.

Equating the forces in Eqs. (1) and (2) and taking components (you do it) we get:

$$
\begin{align*}
m x^{\prime \prime} & =q y^{\prime} B \\
m y^{\prime \prime} & =-q x^{\prime} B  \tag{4}\\
m z^{\prime \prime} & =0
\end{align*}
$$

Forget the third $(z)$ equation since its solution does not couple to those of the $x$ - and $y$-equations. Note that the $x$ - and $y$-equations are "coupled", in that $x^{\prime \prime}$ involves $y^{\prime}$ and $y^{\prime \prime}$ involves $x^{\prime}$. If we define $a(x, y) \equiv(q / m) B(x, y)$ then Eqs. (4) can be written:

$$
\begin{gathered}
x^{\prime \prime}=a y^{\prime} \\
y^{\prime \prime}=-a x^{\prime}
\end{gathered}
$$

or, equivalently,

$$
\begin{equation*}
v_{x}^{\prime}(t)=a(t) v_{y}(t) ; v_{y}^{\prime}(t)=-a(t) v_{x}(t) \tag{5}
\end{equation*}
$$

In the Numerov method we deal with the solution functions $x(t), y(t)$, $v_{x}(t)$, and $v_{y}(t)$, as a series of numbers at "net-point" times that are integrally spaced:

$$
t_{n}=n \Delta
$$

This is illustrated in Fig. 1 for $x(t)$.
We then write:

$$
\begin{aligned}
x_{n} & \equiv x\left(t_{n}\right) \equiv x(n \Delta) \\
y_{n} & \equiv y\left(t_{n}\right) \equiv y(n \Delta) \\
v_{x, n} & \equiv v_{x}\left(t_{n}\right) \equiv v_{x}(n \Delta) \\
v_{y, n} & \equiv v_{y}\left(t_{n}\right) \equiv v_{y}(n \Delta)
\end{aligned}
$$

and our Eqs. (3) and (5) become:

$$
\begin{align*}
x_{n}^{\prime} & =v_{x, n} \\
y_{n}^{\prime} & =v_{y, n} \\
v_{x, n}^{\prime} & =a_{n} v_{y, n}  \tag{6}\\
v_{y, n}^{\prime} & =-a_{n} v_{x, n}
\end{align*}
$$

These are four coupled equations.
We now connect the consecutive values of the $x$ 's and $v$ 's by making Taylor's Series expansions of each of them. For example:

$$
x(t+\Delta)=x(t)+\Delta x^{\prime}(t)+\frac{\Delta^{2}}{2!} x^{\prime \prime}(t)+\ldots
$$

We will choose a sufficiently small so that terms beyond the second will be negligible compared to the first two terms. Then in our net-point notation and using Eq. (6):

$$
\begin{align*}
x_{n+1} & =x_{n}+\Delta v_{x, n} \\
v_{x, n+1} & =v_{x, n}+\Delta a_{n} v_{y, n} \tag{7}
\end{align*}
$$

$\triangleright$ You derive the equations for $y_{n+1}$ and $v_{y, n+1}$.
Given the $t=0$ position and velocity components,

$$
x_{0} ; y_{0} ; v_{x, 0} ; v_{y, 0}
$$

we can use the four "recurrence" relations (7) to generate the four position and velocity components at time $t=\Delta$ :

$$
x_{1} ; y_{1} ; v_{x, 1} ; v_{y, 1}
$$

Putting the latter back into the right hand side of the recurrence relations, we get the values at time $t=2 \Delta$. Continuing this process, we can find
the trajectory as far into the future as we wish. We must only be careful to put the correct value of $\vec{B}$ into $a$ at each space-point $\left(x_{n}, y_{n}\right)$.

Finally, how does one know what size time interval $\Delta$ to use? One could attempt to assess the importance of successive terms in the Taylor's Series, but a more reliable method is to decrease $\Delta$ until the predicted trajectory stabilizes; that is, until it does not change significantly when $\Delta$ is made even smaller. However, one must be aware that if $\Delta$ is continually made even smaller, a point will be reached where the errors will start increasing due to the computer's finite-word-size limit.

The algorithm, then, consists of:

1. recurrence relations
2. method of assuring desired accuracy
3. insertion of initial conditions.

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